

The background of the slide is a composite image. The top half shows a night sky filled with stars and a vibrant aurora borealis (Northern Lights) in shades of green, yellow, and orange. The bottom half shows a dark, snowy landscape with some rocky outcrops. A large, semi-transparent teal shape covers the bottom half of the image, serving as a backdrop for the text.

The Geometry of Propagation

Heatherly Takeuchi – N6HKT

Agenda

Introduction

Presentation by José Nunes CT1BOH (48:47)

MUF Isolines

Geometry of Earth

Bringing it together





Introduction

Propagation is a mystery especially when working HF

Is it predictable? More so every day.



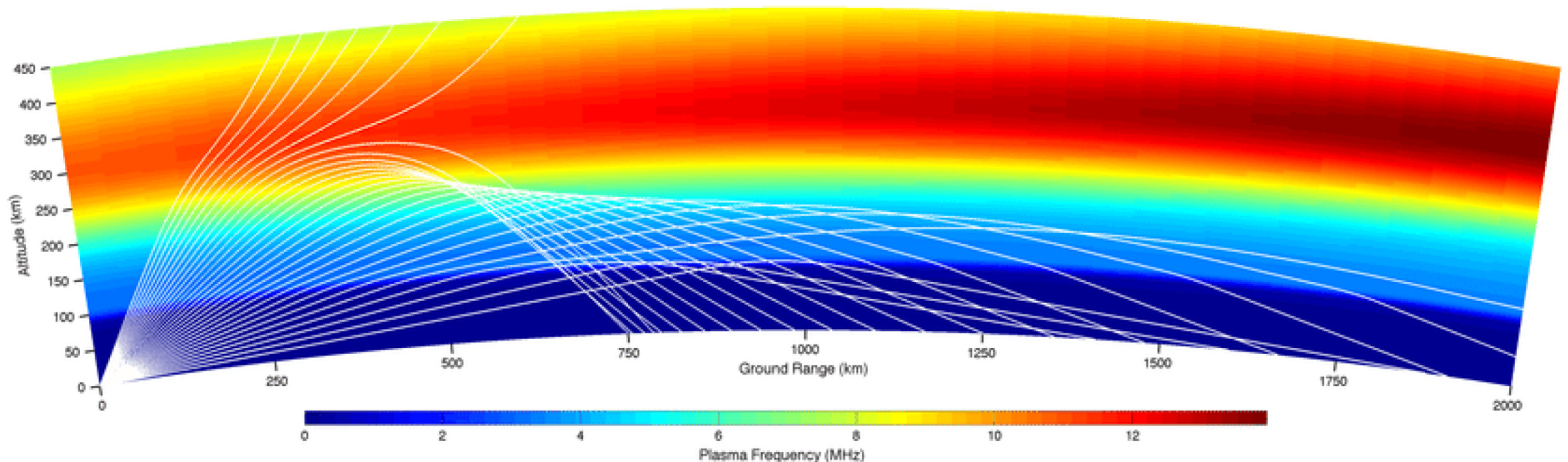
Opening Thoughts

- This is not a talk about the Solar cycle, the magnetosphere, or the ionosphere.
- Just because you can hear them does not mean they can hear you and vice versa.

There is nothing magic about propagation

In search of MUF isolines

Contest University

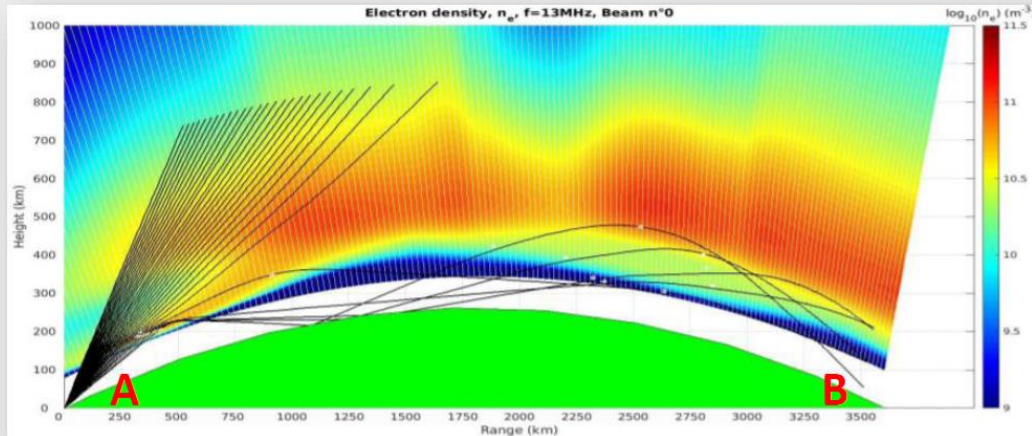


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ct1boh@gmail.com

Maximum Usable Frequency (MUF)

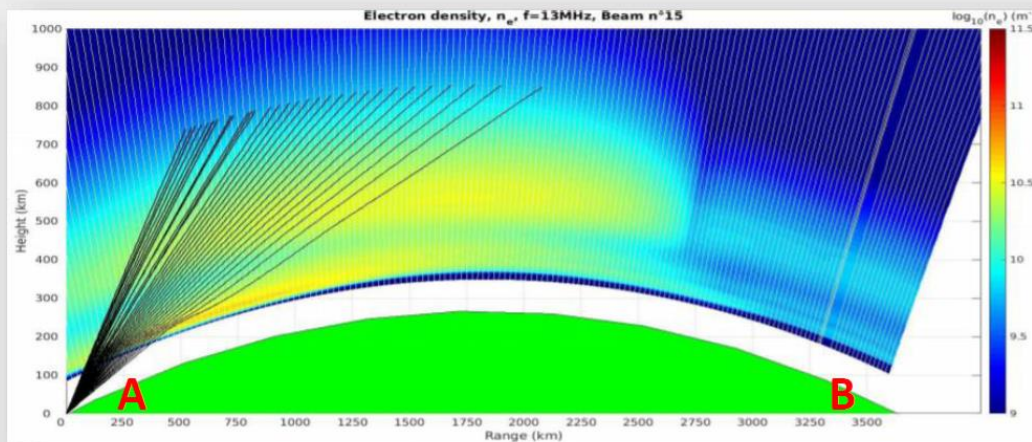
- In radio transmissions, MUF is the highest radio frequency that can be used for transmission between two points via reflection (refraction) from the ionosphere (sky-wave propagation) at a specific time, independent of transmitter power.

Sky-wave HF communication from A to B relies on ionosphere for refraction, otherwise signal will be lost into space



Refraction of signal – There is propagation

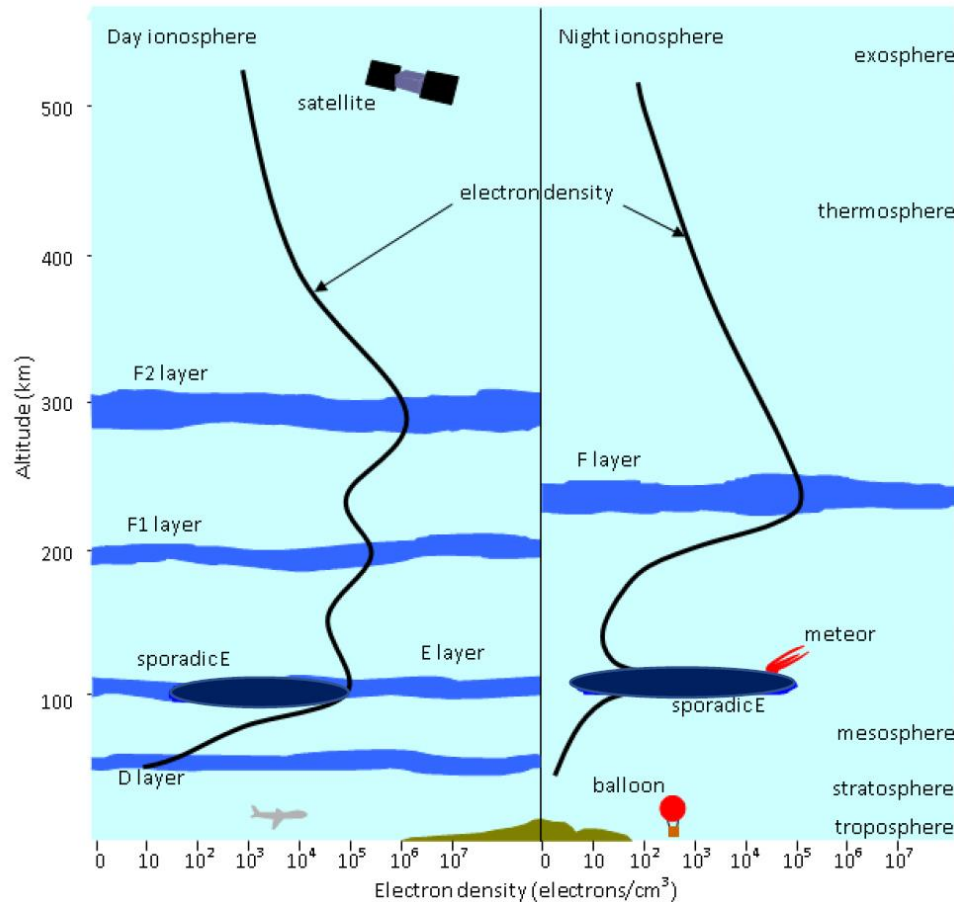
- Electron density of refraction area
- D layer absorption level
- E, F1 and F2 refraction layers
- MUF in refraction area
- Operating frequency below critical frequency in refraction area (or “MUF of ionosphere”)



No refraction of signal – no propagation

- No refraction in one of the ionospheric layers
- Operating frequency above MUF of E, F1 or F2 layer
- High angle of incidence of RF rays
- Dead band or no propagation to a particular area of the world

Different Layers of the ionosphere play a pivotal role in absorption and refractions of radio signals together with critical frequency



Layers

- Layers to refract signals if MUF is higher than operating frequency
- Refractions depends on the density of electrons
- Density depend on ionization from the sun (dominant) and coupling with troposphere
- Electron density with peaks and inflection points (layers)
- The layer with the highest density determines MUF

D layer

- D layer has no ability to refract signals
- D layer absorbs signals passing through it reducing signal strength
- Absorption levels can vary with season and sun activity (x-rays)
- D layer is not present during the night

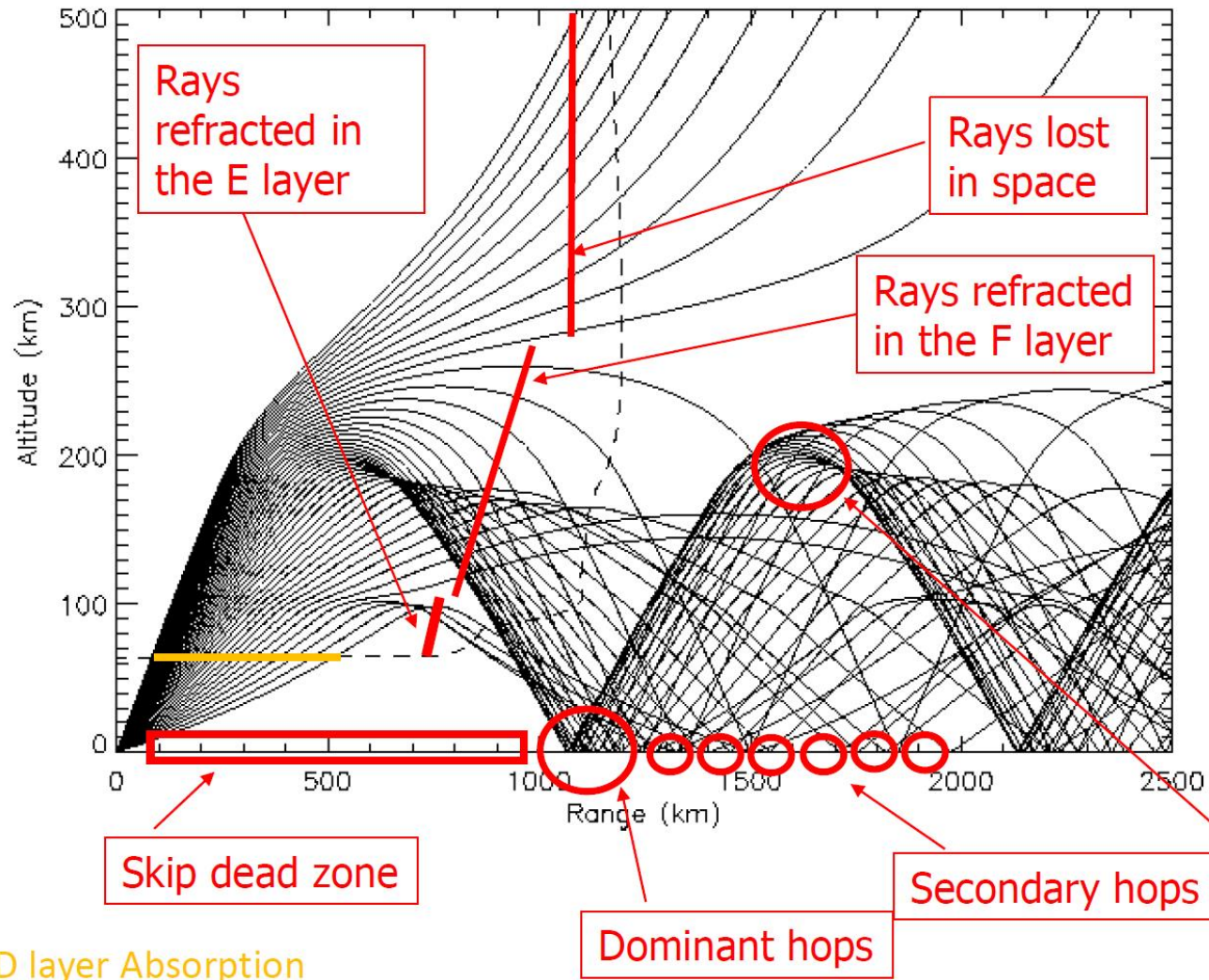
E layer

- E layer is the first layer capable of refracting signals
- It is present day and night
- Sporadic E occurs with intense electron density
- When E layer has the highest electron density it blankets signals into F layer

F1, F2 layer

- F2 is the most reliable layer for long distance communications
- At night F1 and F2 layers merge into F
- Being the highest layer, it requires less hops in the refractions
- Usually, the MUF is provided by the F layer

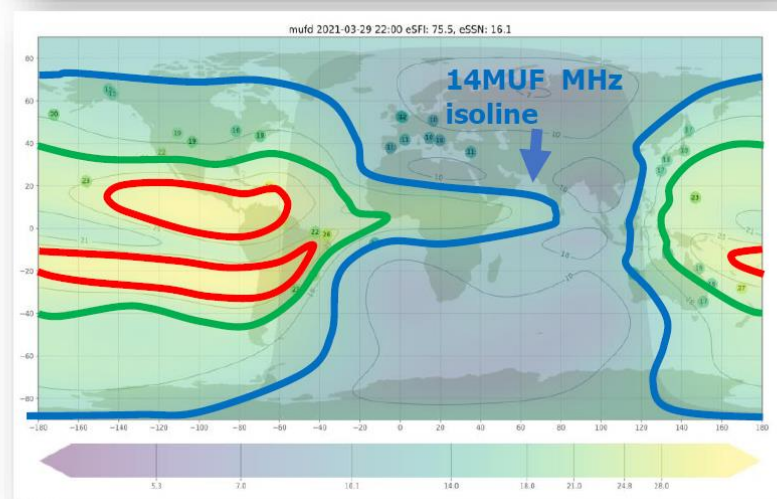
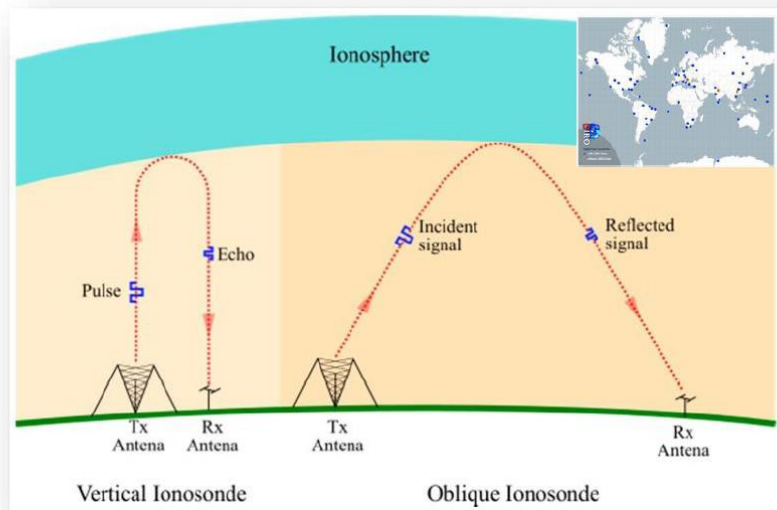
Understanding signal refraction, skip distance, multi-hop and critical frequency ("MUF of ionosphere")



Critical frequency (MUF of ionosphere)

- For propagation to exist, RF signals must refract in the ionosphere back into earth
- Refraction will depend on the angle of incidence of the signal ray into the ionosphere
- Refraction will happen in ionosphere if the critical frequency of ionosphere is greater than the operating frequency
- If the "MUF" of ionosphere in the refraction point is lower than the operating frequency, signal will be lost in space

Probing the ionosphere with a limited network of Ionosondes enables to model ionosphere real-time conditions



A worldwide network of ionosondes

- The ionosonde is used to find the NIVS MUF frequency in the ionosphere immediately above the antenna
- The result of the probe of the ionosonde is an ionogram, depicting ionospheric layers, the height of each layer and the electron density
- A model enables to transform NVIS MUF into MUF for 3000Km
- Oblique ionosonde depicts the true MUF between two points

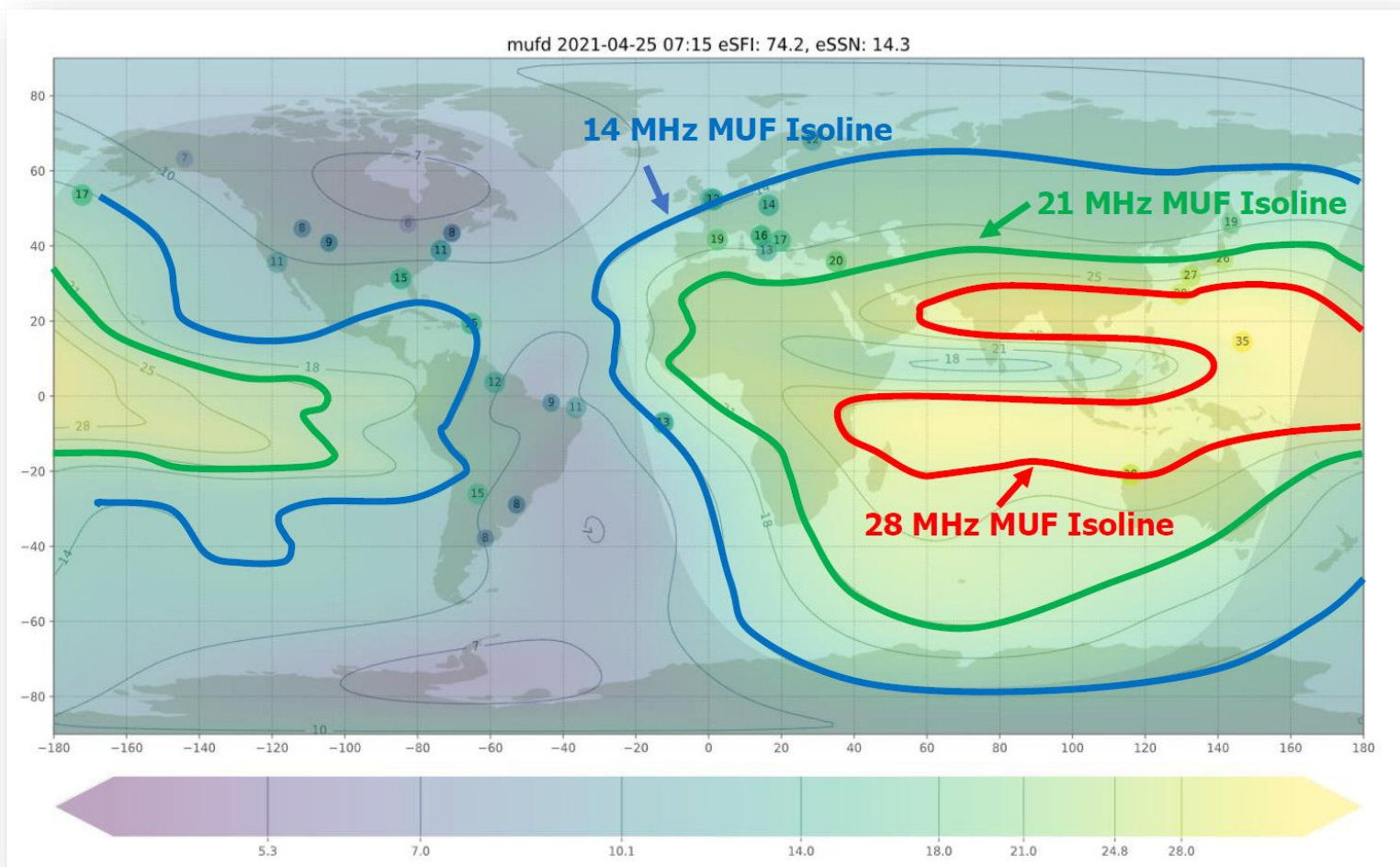
Worldwide MUF isoline map model

- Worldwide isoline MUF map shows areas of the world that support a signal to be refracted
- The IRI (International Reference Ionosphere) model uses data from the network of ionosondes around the world, compiled by NOAA and GIRO (Lowell Global Ionospheric Radio Observatory)
- Most ionosondes are clustered around few areas of the world, with large areas away from a probe

<https://www.digisonde.com/>

<https://prop.kc2g.com/>

What is a MUF isoline and why the MUF isoline limits propagation to refraction points inside the MUF isoline

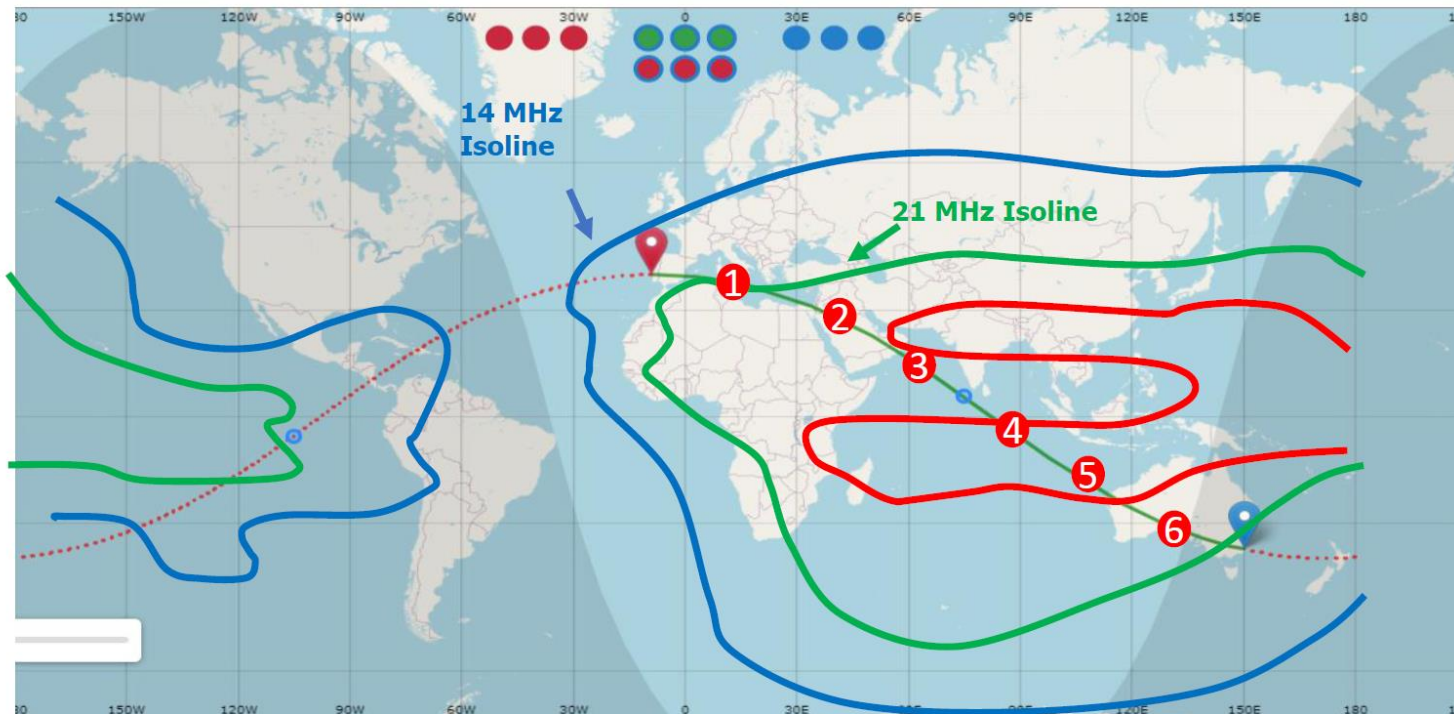


MUF isoline sets propagation

- A MUF isoline is a contour line where the critical frequency (or ionosphere MUF) is the same all over the world
- Inside the 14 MHz isoline MUF frequency is higher than 14 MHz and outside it is lower
- Propagation on 14 MHz will only be possible if the refraction points in the ionosphere are inside the area of the MUF isoline limit

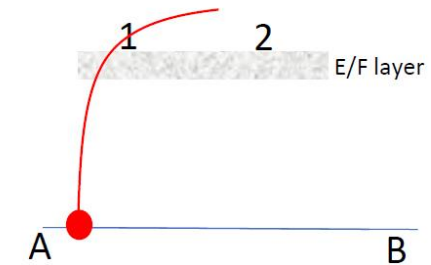
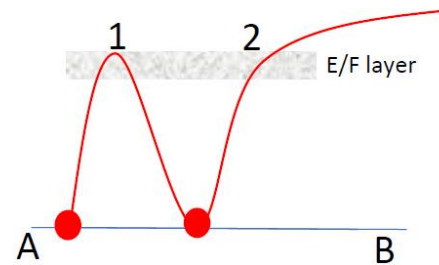
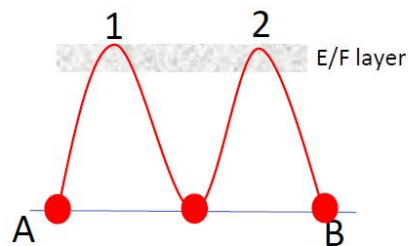
<https://prop.kc2g.com/>

Paths that have refraction points inside or near the edge of the MUF isoline will hold propagation

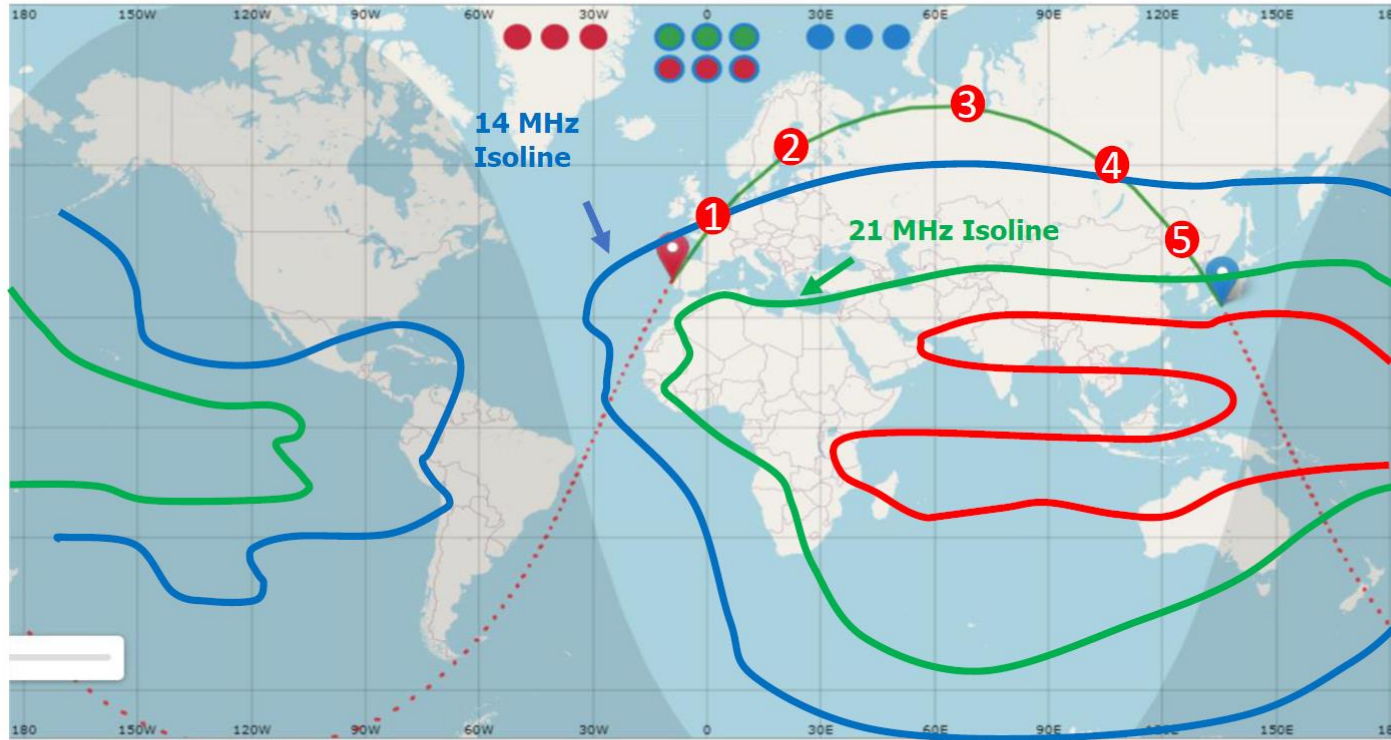


Portugal into Australia path

- The path from Portugal into Australia has all refraction points inside the MUF limit isoline for 21 MHz and 14 MHz
- All refraction points will find a MUF frequency above the operating frequency (21 MHz)
- The only limiting factor is power, gain and D layer absorption

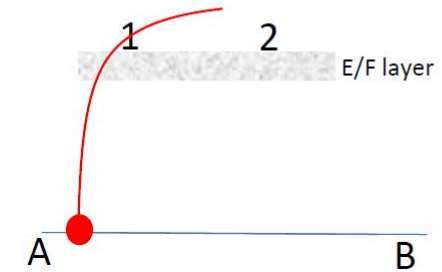
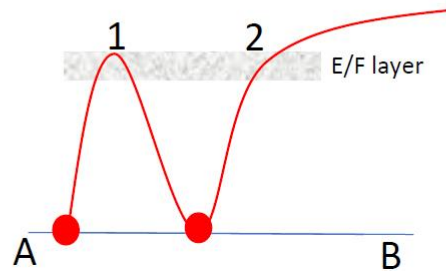
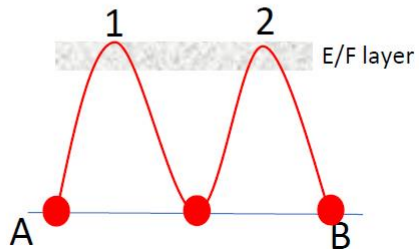


Paths that have refraction points outside of the MUF isoline will not hold propagation



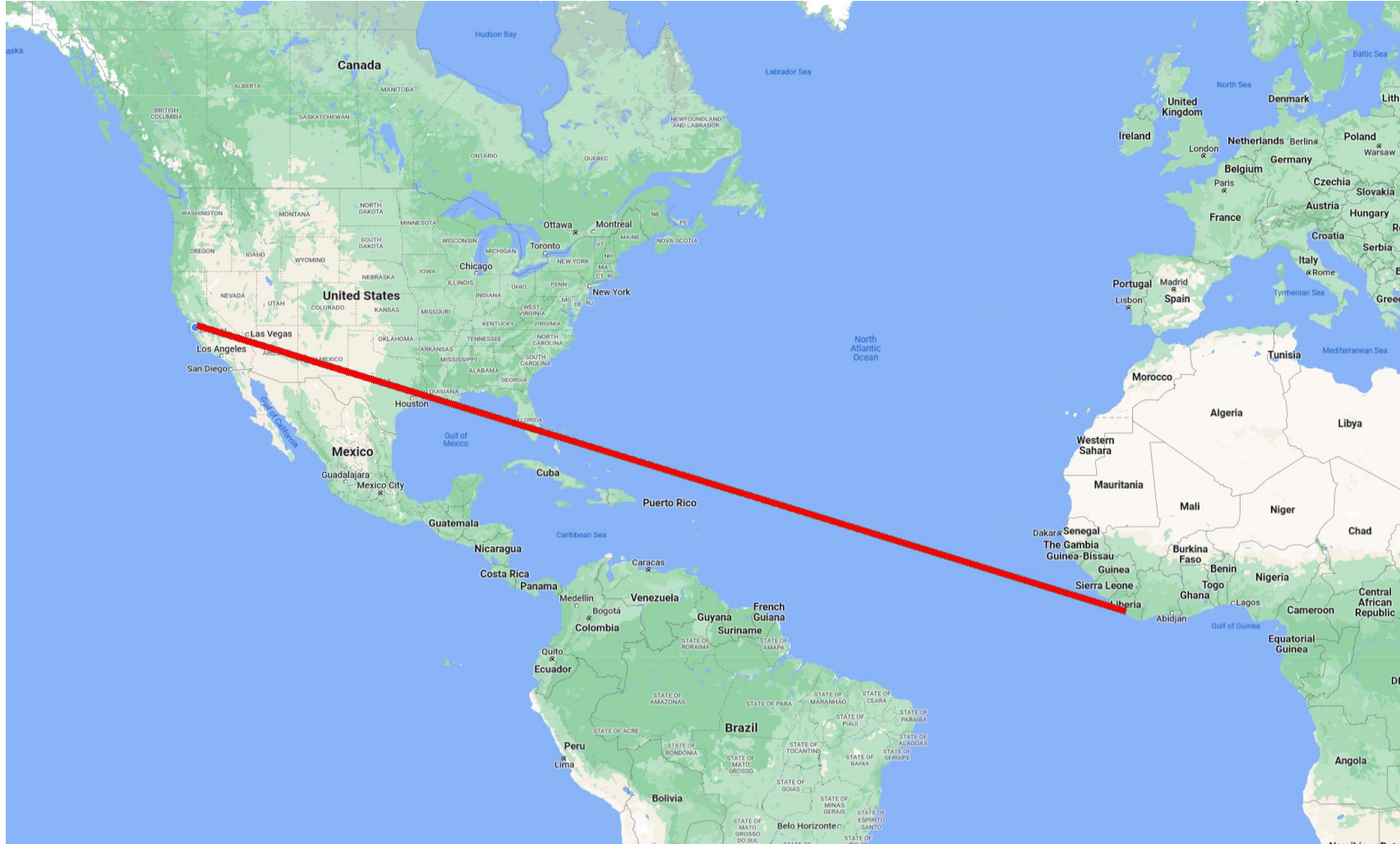
Portugal into Japan path

- The path from Portugal into Japan has refraction points outside the MUF limit isoline
- Refraction points 2 and 3 have a MUF frequency lower the operating frequency (14 MHz)





A straight line is the shortest distance...



Flat map of Earth





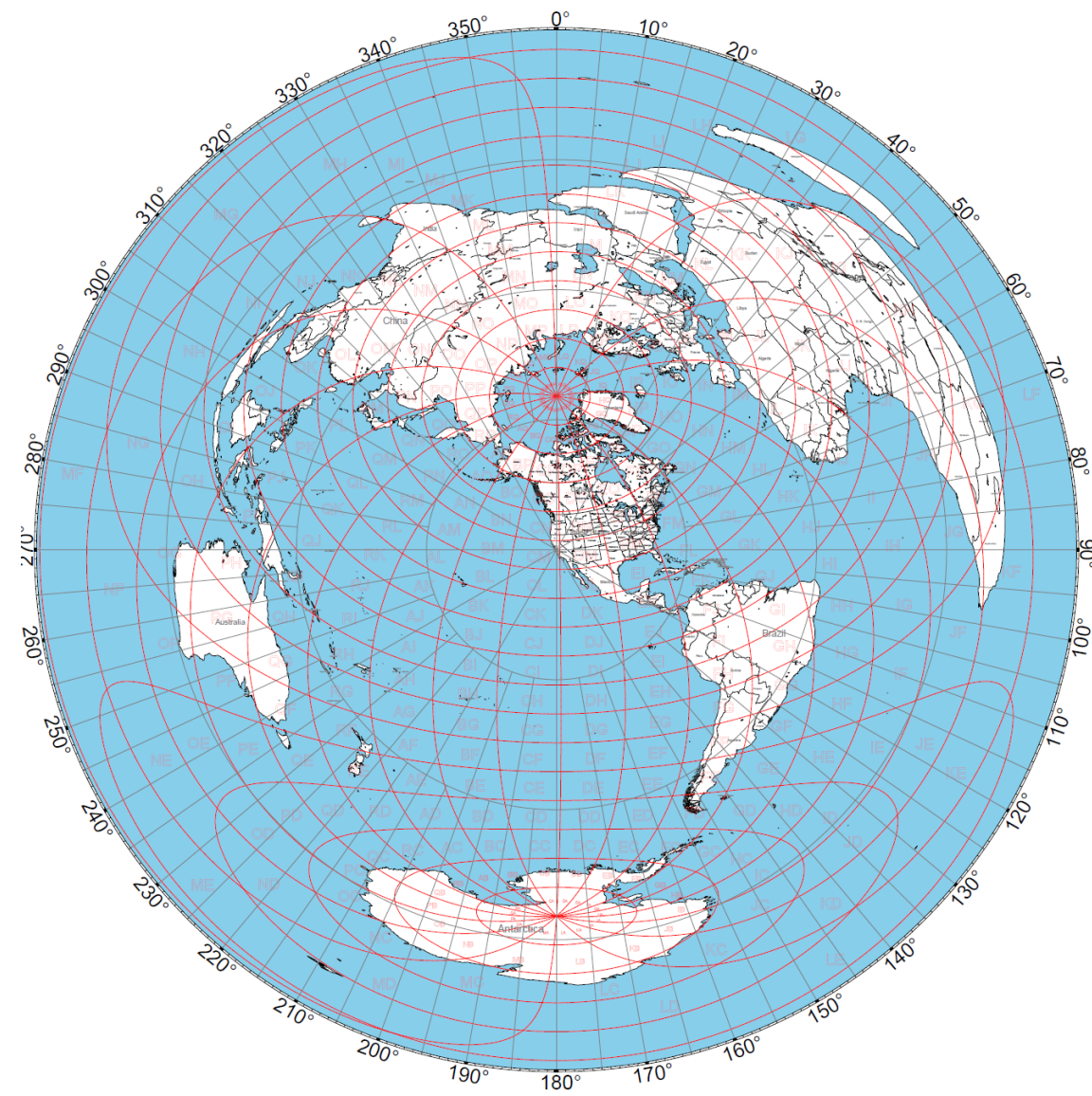
How to See The World for HF



Azimuthal Map

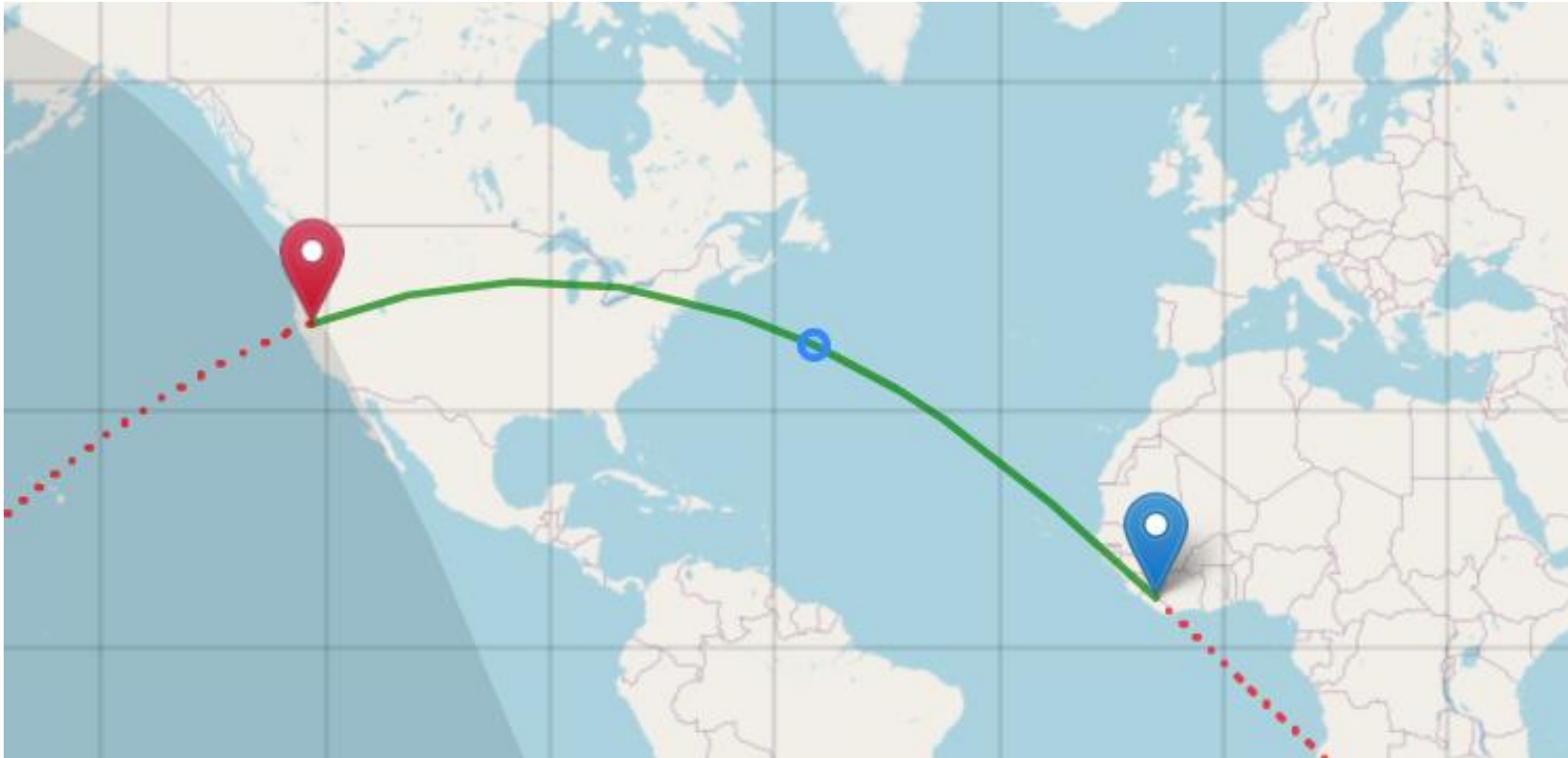
Center: 36°49'59"N 121°25'0"W

Courtesy of Tom (NS6T)



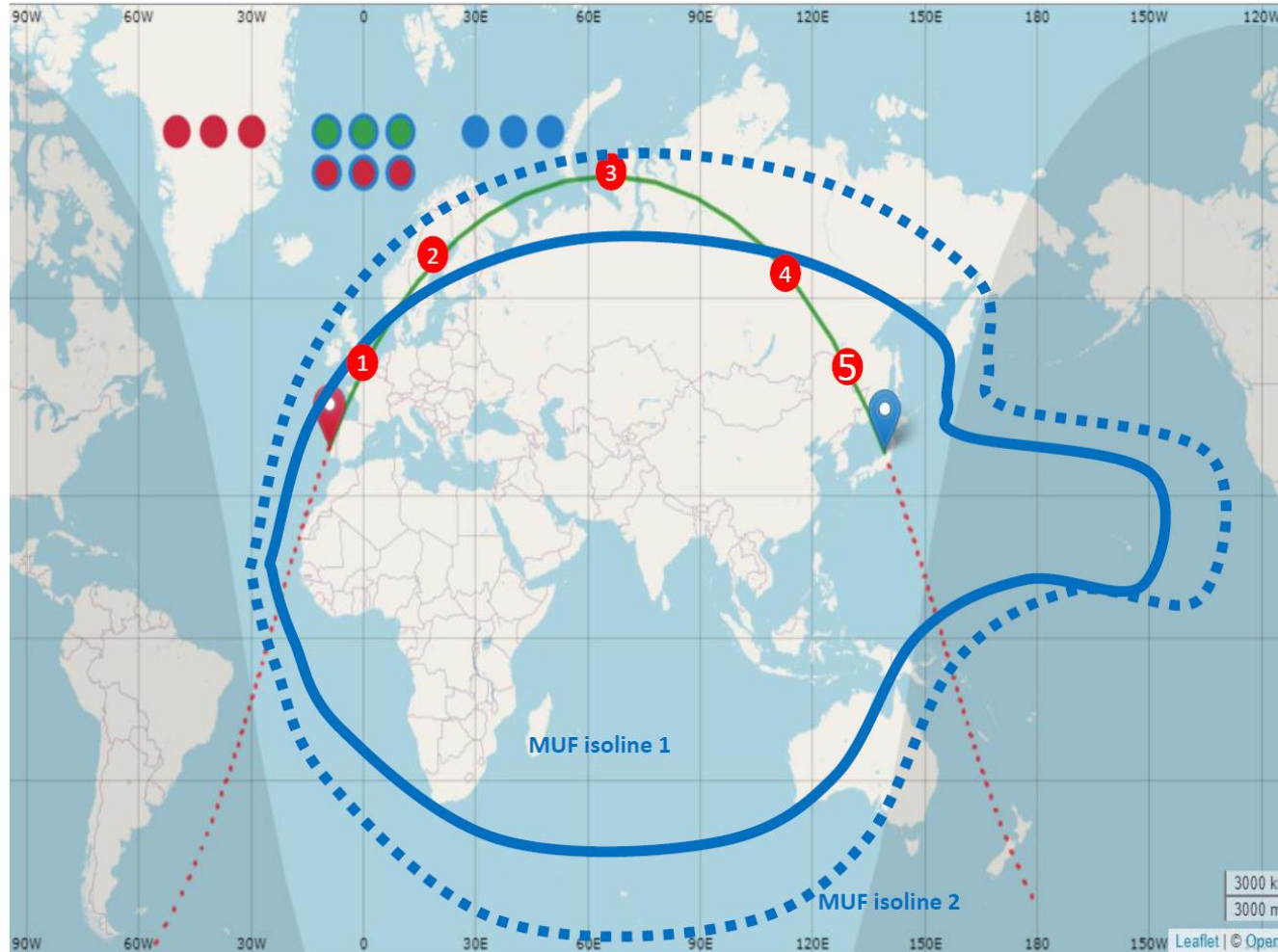
How to See The World for HF

How to See The World for HF



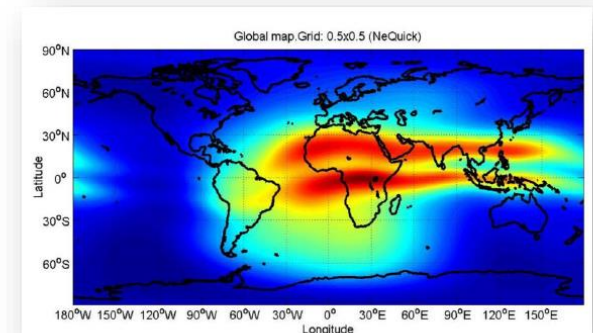
- <https://www.voacap.com/hf/>

Variability in propagation can be seen and the increase and decrease of the shape and size of the MUF isoline figure



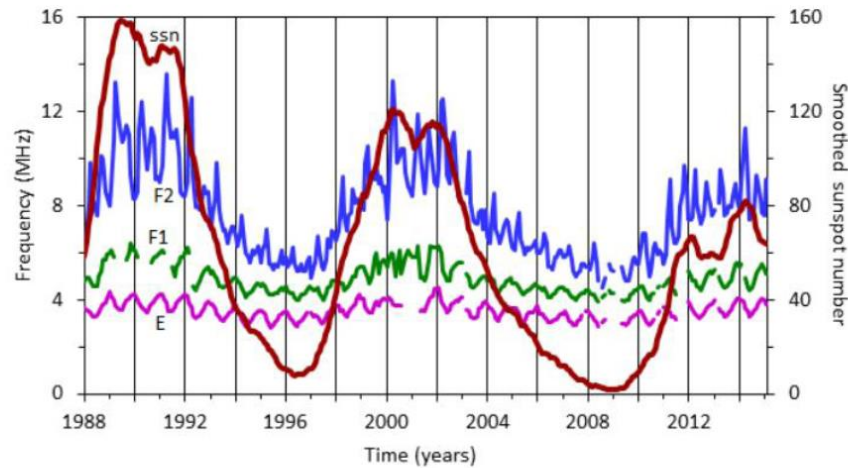
The shape and size of the MUF isoline

- The shape and size of the MUF isoline determining the quality of propagation is explained by a multitude of factors:
 - Solar Cycle evolution
 - Seasonal variations
 - Latitude variations
 - Diurnal variations
 - Solar and tropospheric events
 - Layer path mode variability
 - Magnetosphere



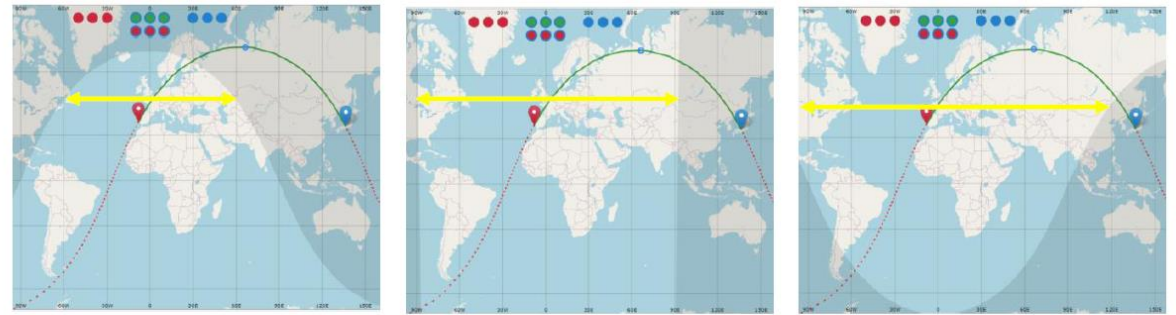
Ionospheric variability – solar cycle and seasonal variations

Solar cycle evolution



- Increased radiation during the peak of solar cycle, increases electron density and F2 layer MUF and to a less extent F1 MUF
- Sun radiation does not have a noticeable impact in E layer MUF

Seasonal variations



Winter Solstice

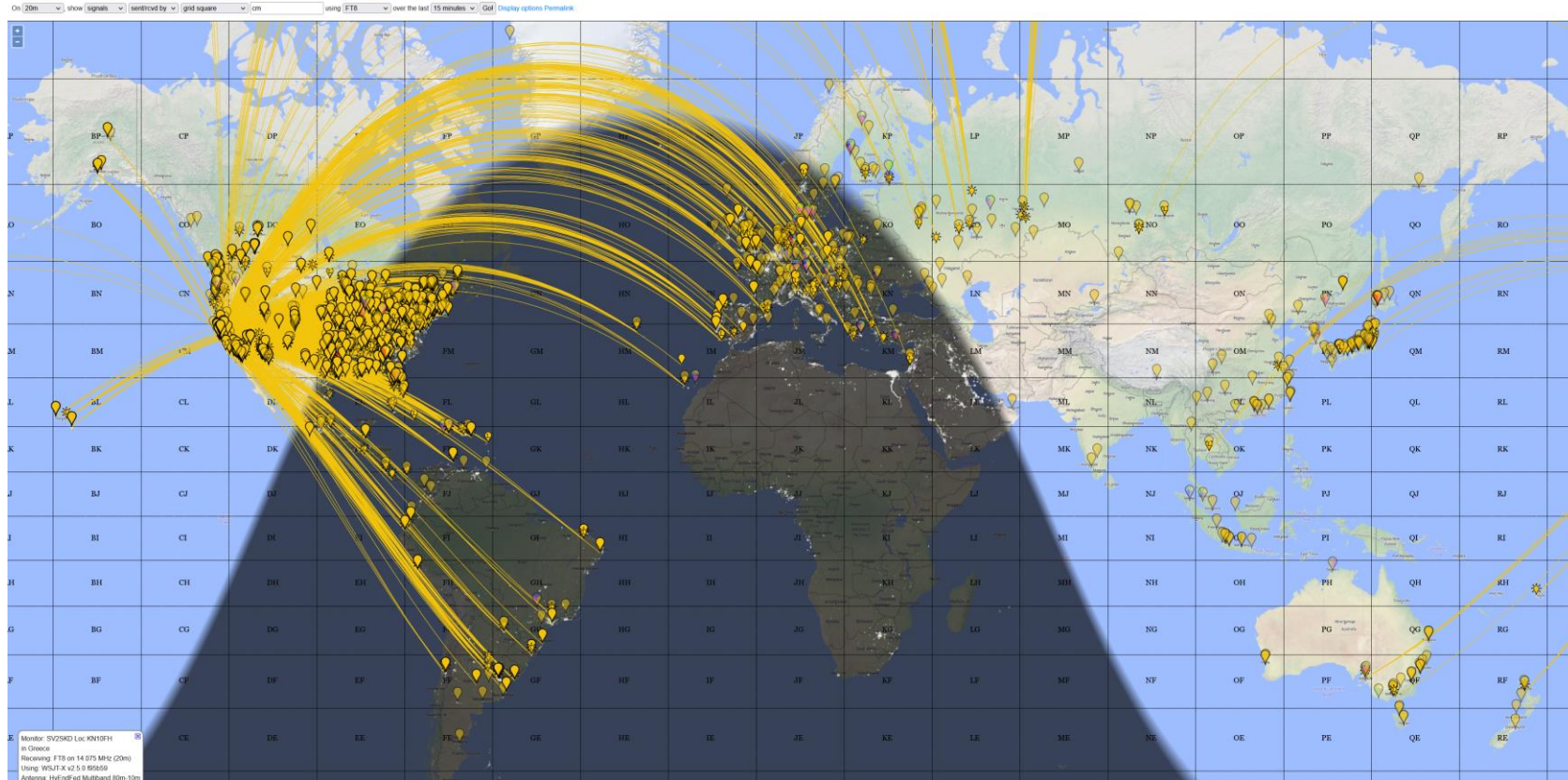
Equinox

Summer Solstice

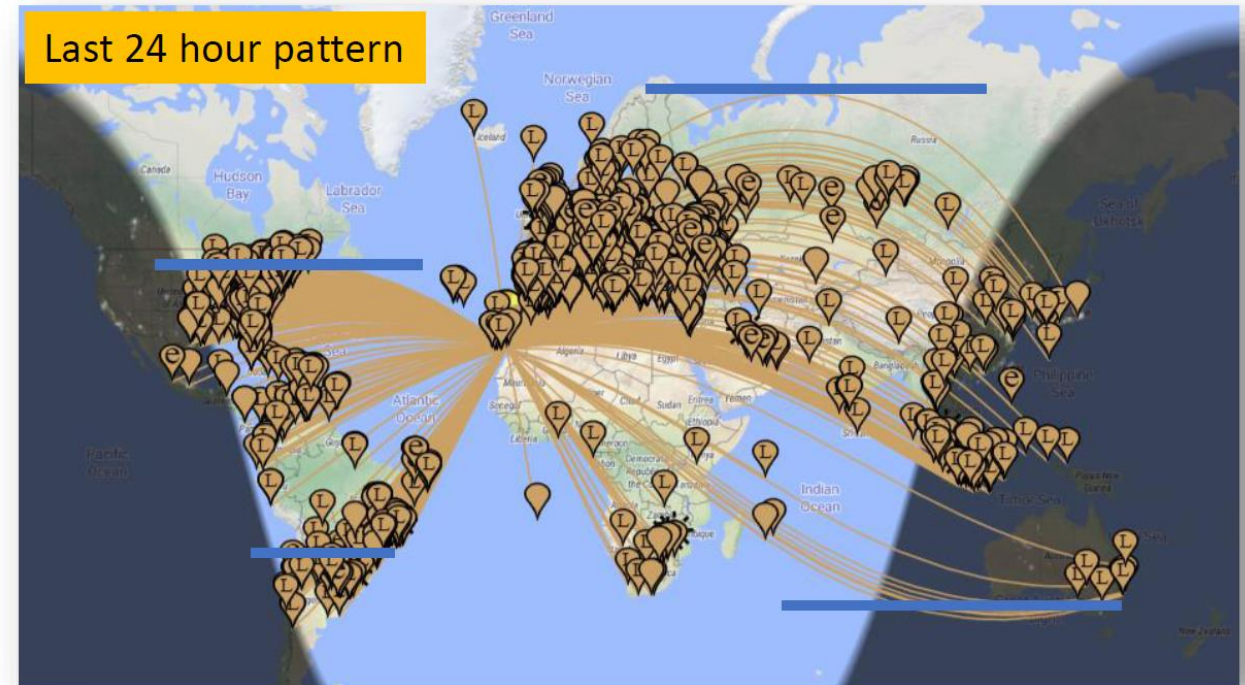
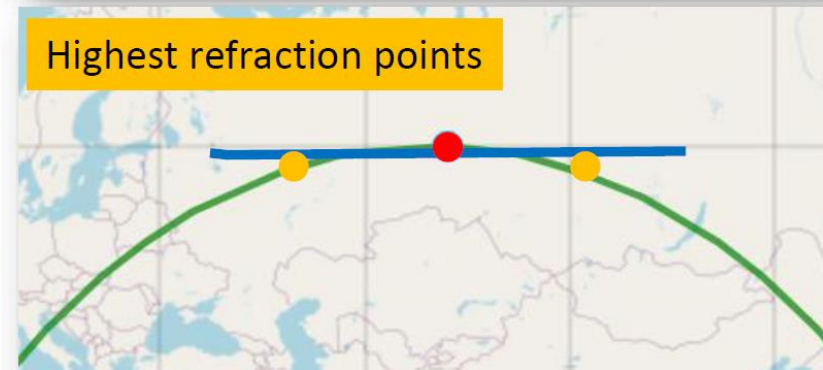
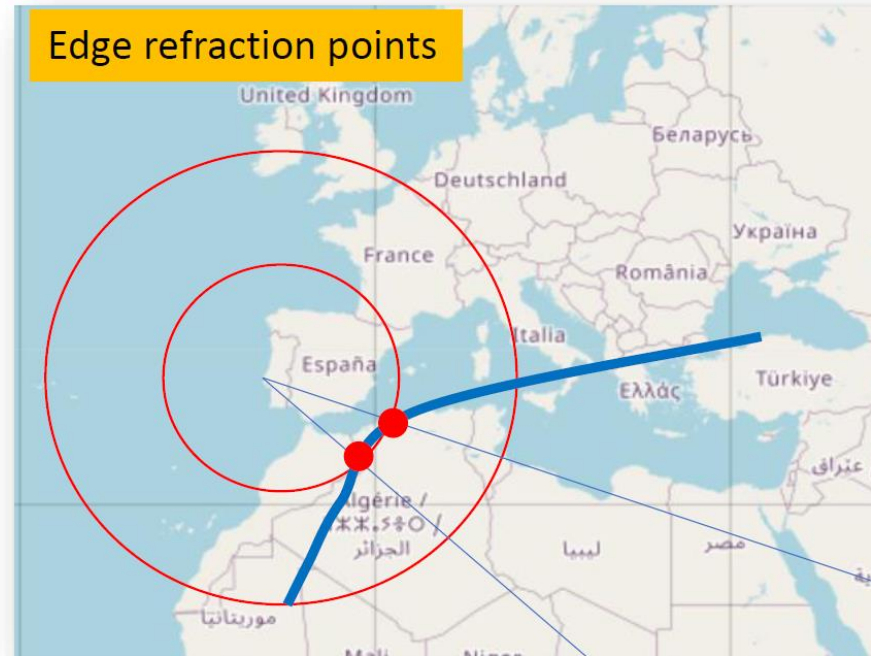
- Seasonal variations have a profound impact in ionization of atmosphere
- Higher or lesser day hours provide different hours of illumination and ionization that impact MUF
- Increased number of refraction points in illuminated atmosphere provide better opportunities between longer paths in HF

PSKReporter.info

- Go to: <https://pskreporter.info/pskmap.html>



In search of MUF isolines – looking for refraction points in the edges and in the highest latitudes of the path and getting a 24 hour view for propagation pattern



The exercise of connection the dots...

- Looking for first refraction points into the edges
- Getting the MUF isoline size and shape by looking at last 24-hour spots pattern
- Looking for the highest latitude refraction points

Closing thoughts for the Geometry of Propagation

- Get an Elmer/Elma
- Use the tools
- Ask questions
- Get on the Air!



A vibrant image of the Aurora Borealis (Northern Lights) in shades of green and blue, set against a dark, starry night sky. The aurora's light patterns are visible on the left side of the slide, partially overlapping a teal-colored circular graphic.

Thank You

73,

Heatherly Takeuchi/N6HKT

<https://sbcara.org>